

# Measuring Geo-Neutrinos



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# Earth

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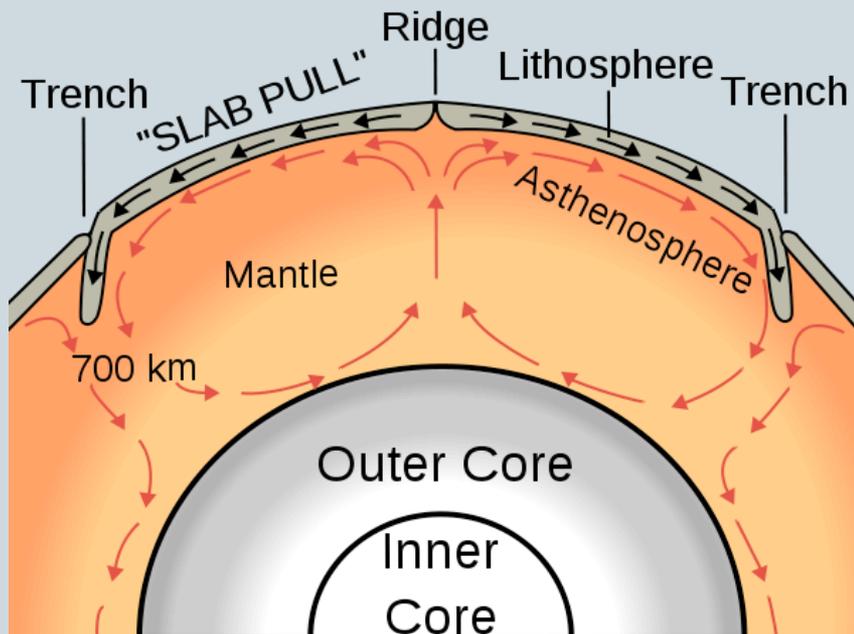


Image: by Surachit, [http://en.wikipedia.org/wiki/File:Oceanic\\_spreading.svg](http://en.wikipedia.org/wiki/File:Oceanic_spreading.svg)

- Seismic data splits Earth into 5 basic regions:
  - inner core
  - outer core
  - mantle
  - oceanic crust
  - continental crust
- All these regions are solid except the outer core.
- The mantle convects even though it is solid.
- Oceanic crust is being renewed at mid-ocean ridges and recycled at trenches.

# Radioactivity in the earth

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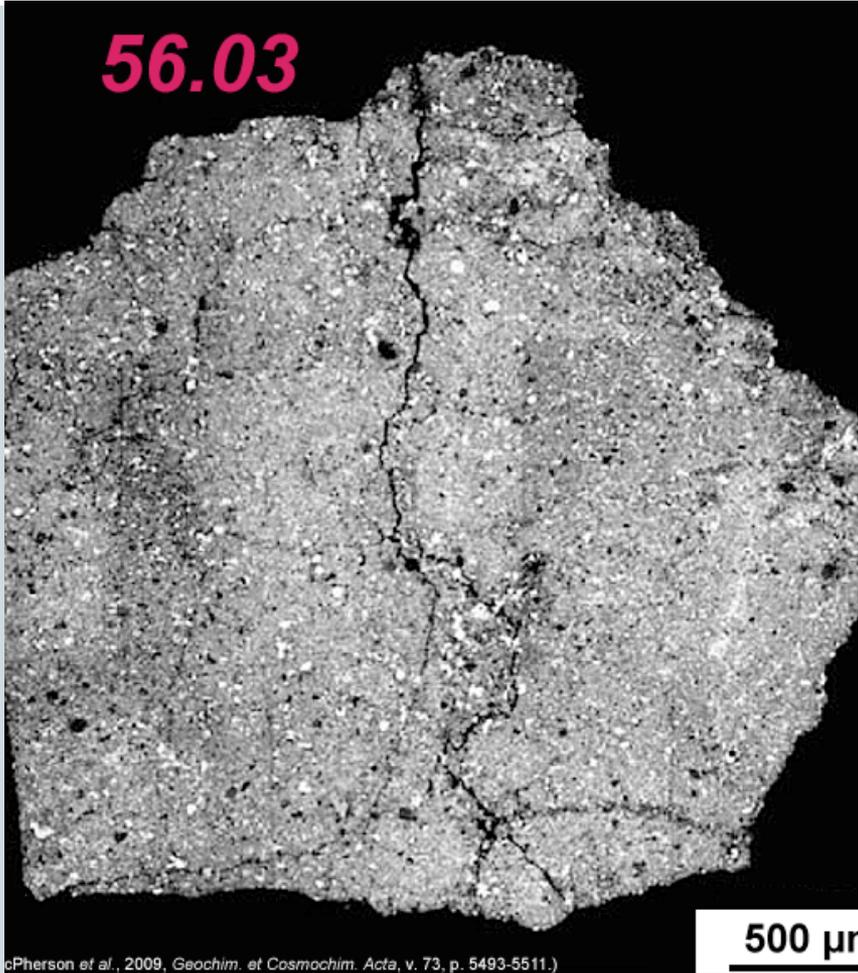


Image: <http://www.psr.d.hawaii.edu/>

- Based on carbonaceous chondrites Lyubetskaya & Korenaga (2007) predict the following concentrations of radioactive isotopes in the earth's primitive mantle and crust
  - U ( $17 \pm 3$ ) ng/g
  - Th ( $63 \pm 11$ ) ng/g
  - K ( $23 \pm 5$ )  $\mu$ g/g
- The resulting heat production rate is  $(16 \pm 3) \times 10^{12}$  W

# Heat flow

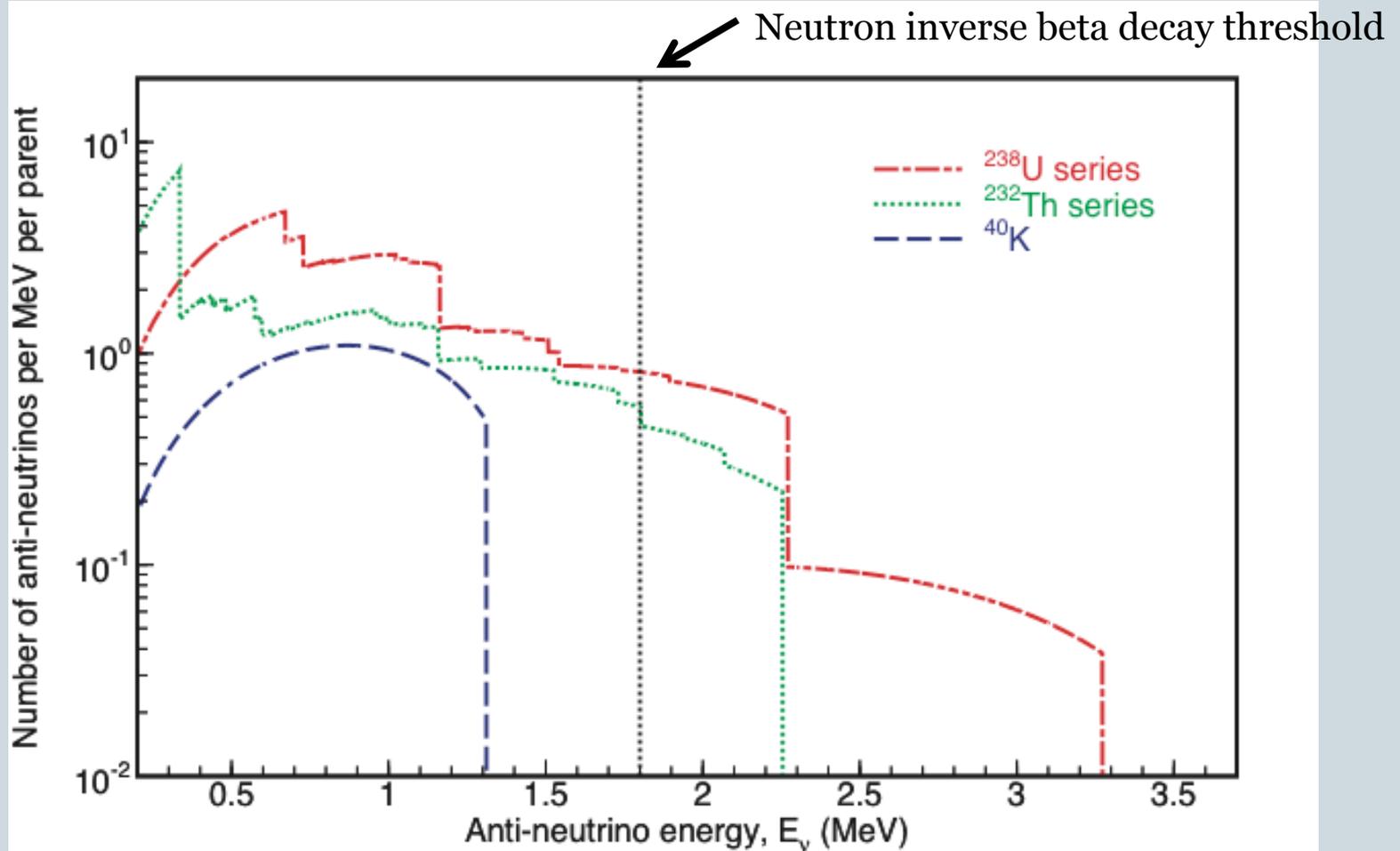
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Total heat loss	$46 \pm 3$ TW
Total radiogenic heat production	$16 \pm 3$ TW
Crust + lithosphere heat production	4.9-8.8 TW
Present Urey ratio	0.11 - 0.34

- Mantle convection models suggest Urey ratios (ratio of heat production to dissipation in the mantle) greater than 0.7
- Problem with
  - Mantle convection model?
  - Total heat loss measured?
  - Estimated radiogenic heat production rate?
- Geo-neutrinos can cross-check the radiogenic heat production.

# Geo-neutrino energy spectrum

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# Where do the neutrinos come from?

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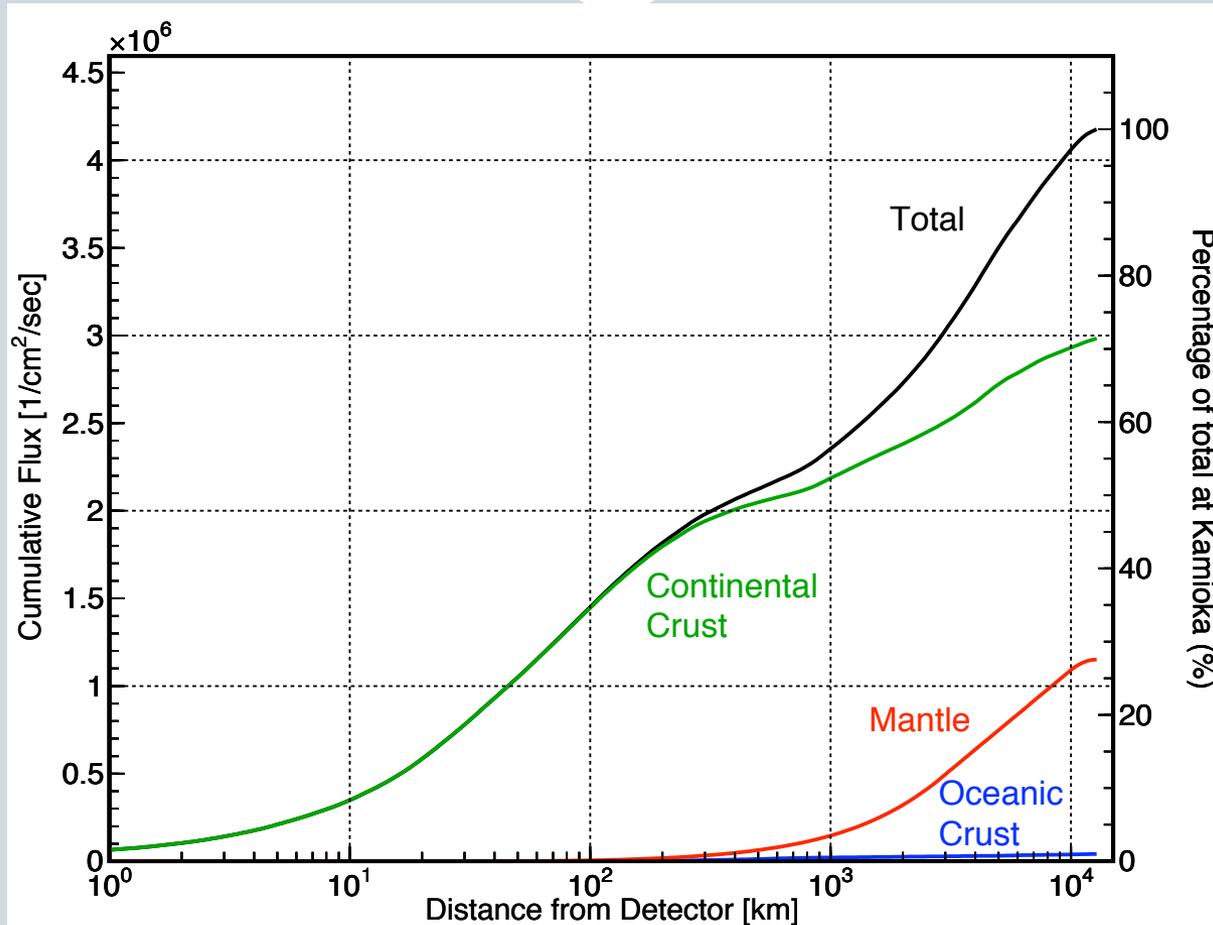


Image: S. Enomoto

# Signal-to-noise ratio

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S/N Ratio: Mantle / (Crust + Reactor)

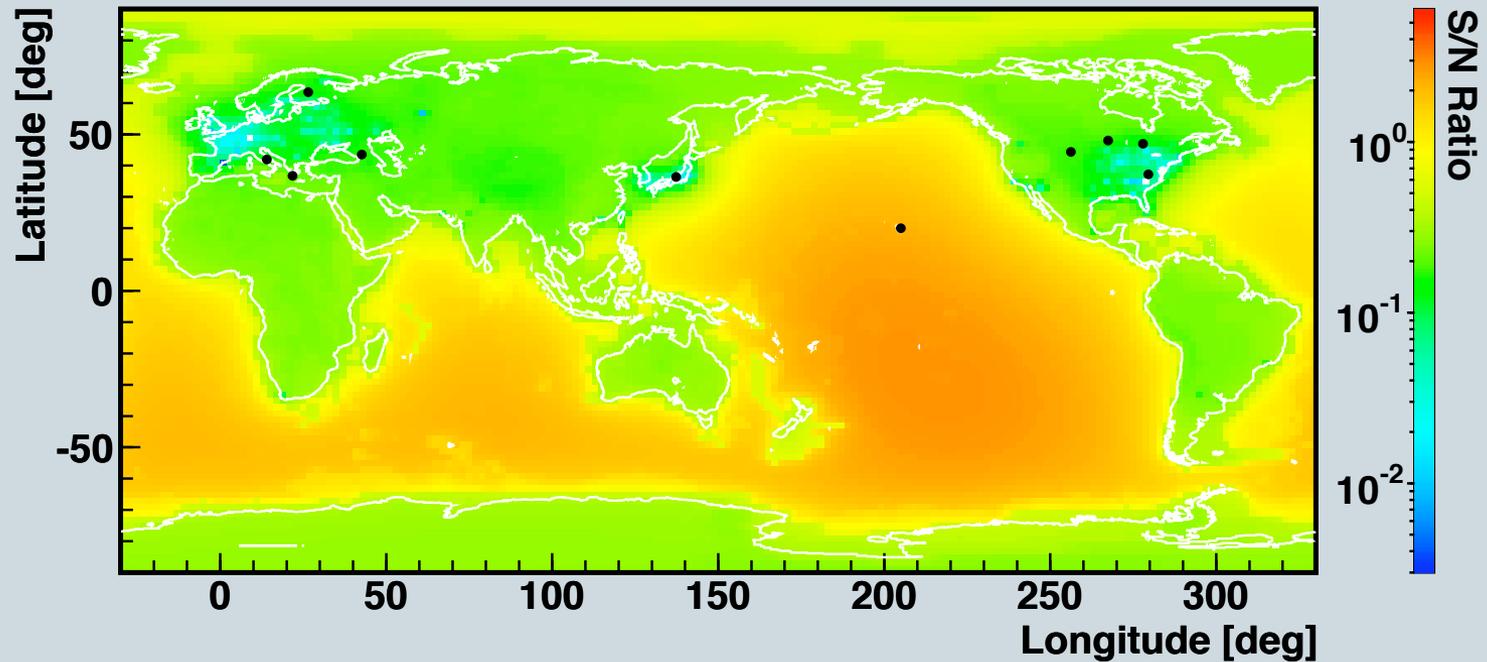


Image: S. Enomoto

# Geo-neutrino experiments

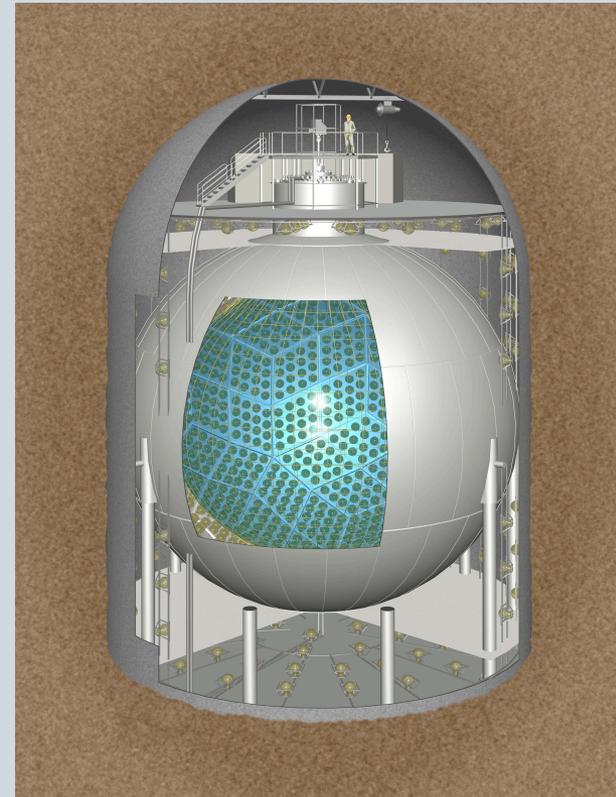
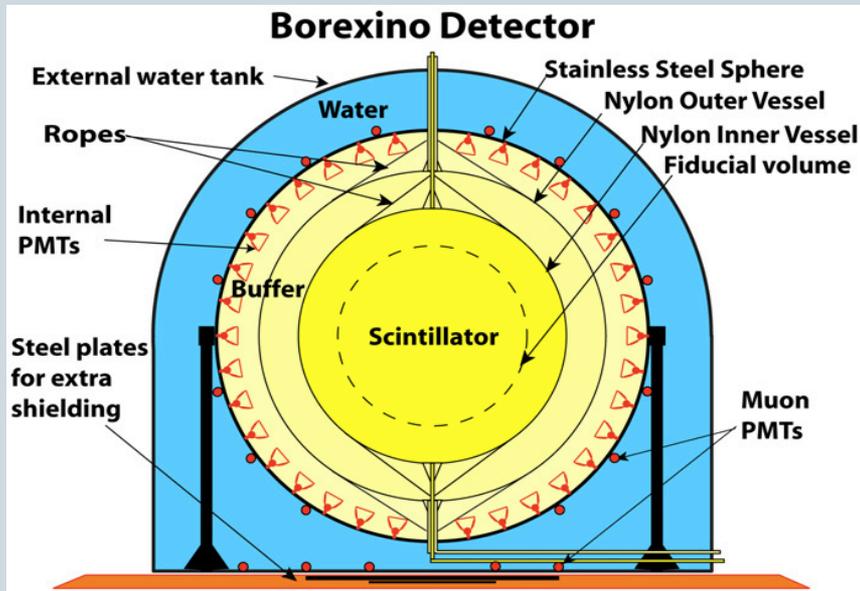
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# Experiments

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**Borexino**  
300 ton, Italy

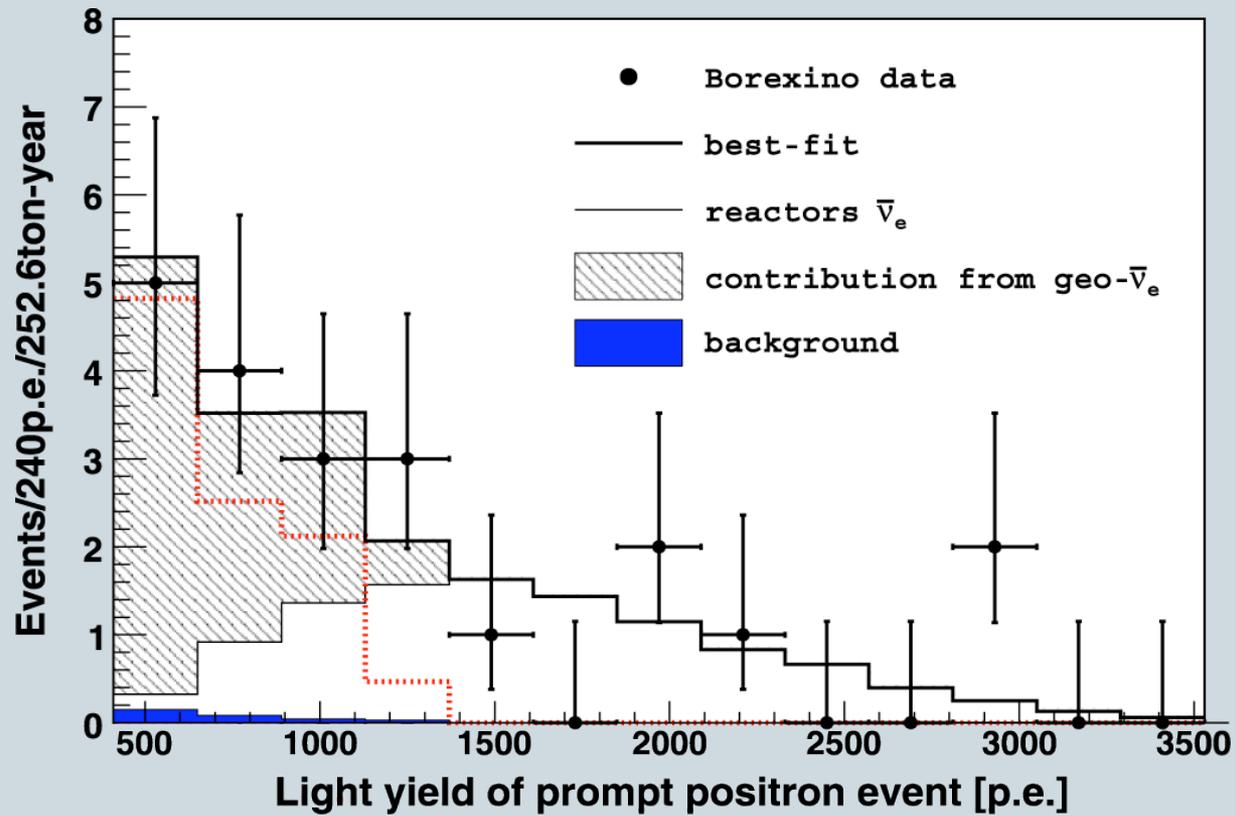
**KamLAND**  
1000 ton, Japan



# Borexino results

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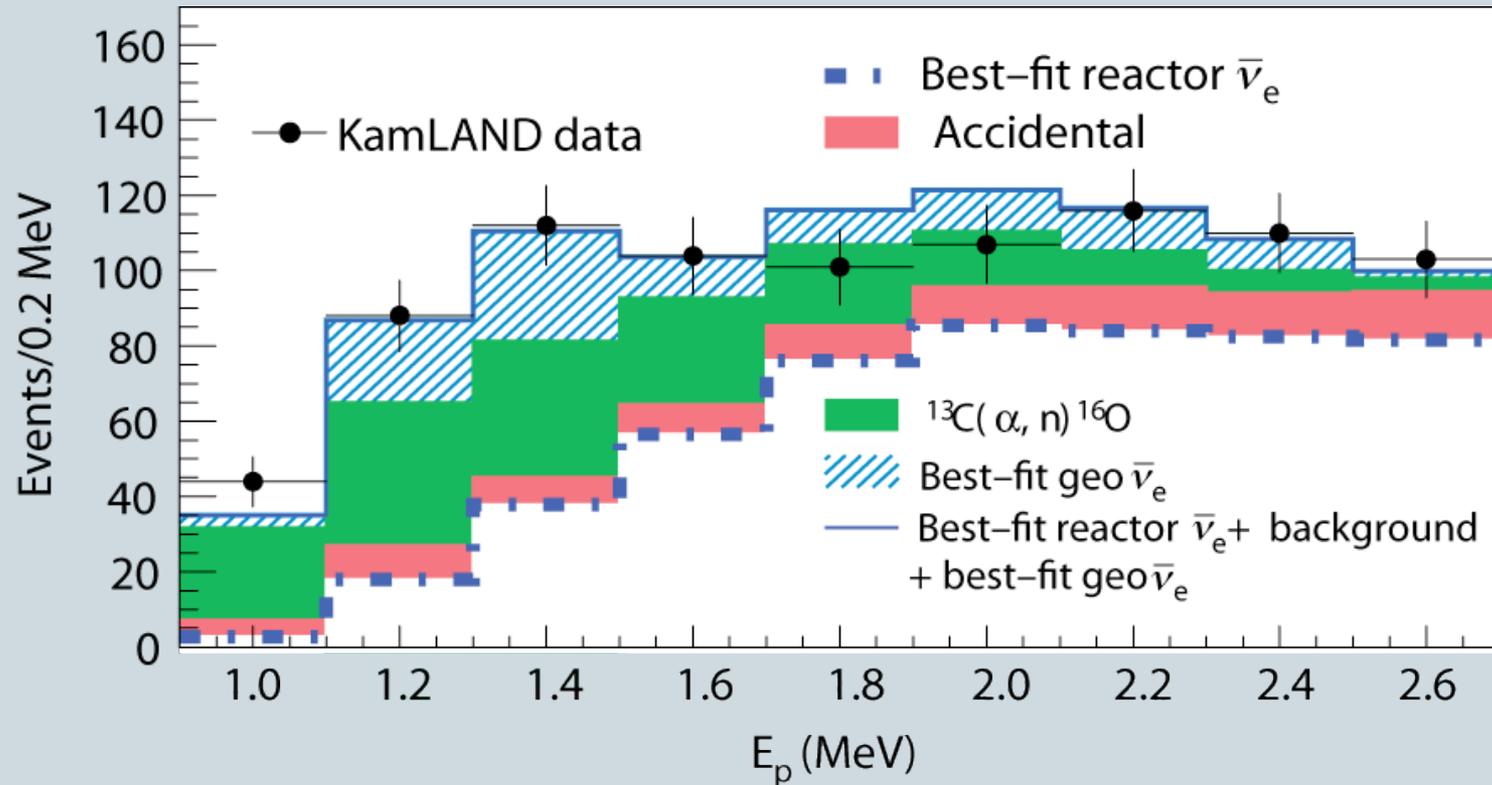
$\sim 0.25 \times 10^{32}$  proton.years



# KamLAND results

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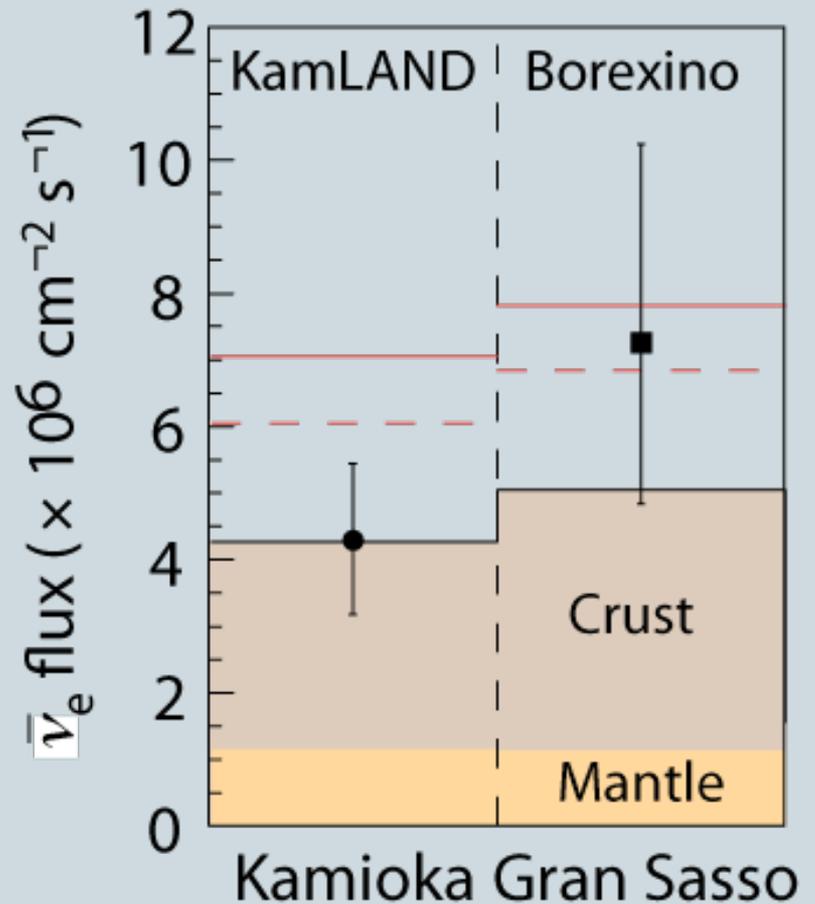
$3.5 \times 10^{32}$  proton.years



# Comparison to predictions

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	KamLAND [ $\times 10^6$ /cm <sup>2</sup> / sec]	Borexino [ $\times 10^6$ /cm <sup>2</sup> / sec]
S. Enomoto <i>et al.</i> (Total)	4.4	5.2
(Crust)	3.2	4.0
(Mantle)	1.2	1.2
F. Mantovani <i>et al.</i> (Total)	4.0	4.6
Measured	4.3 <sup>+1.2</sup> <sub>-1.1</sub>	7.1 <sup>+2.9</sup> <sub>-2.4</sub>



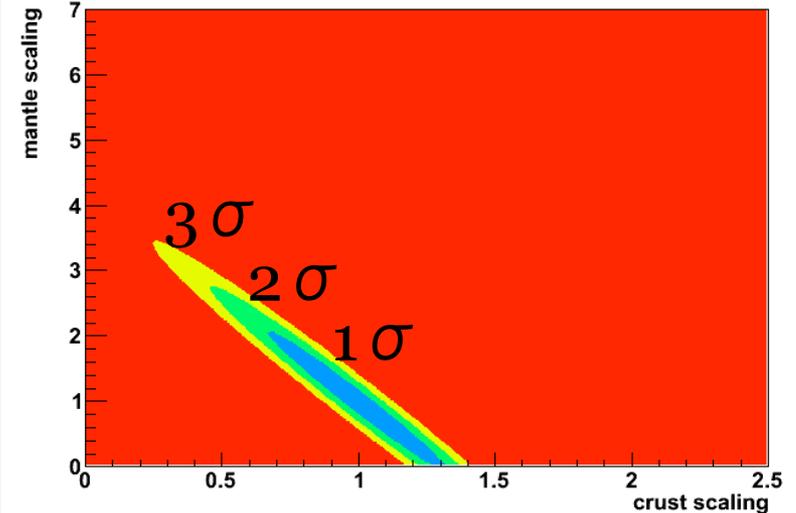
Nature Geoscience **4** 648, 2011

# Need for more detectors

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- Ideally measurement in the ocean to probe the contribution from the mantle.
- Multiple measurements to probe variations.

5% measurement at KamLAND, Borexino, and SNO+

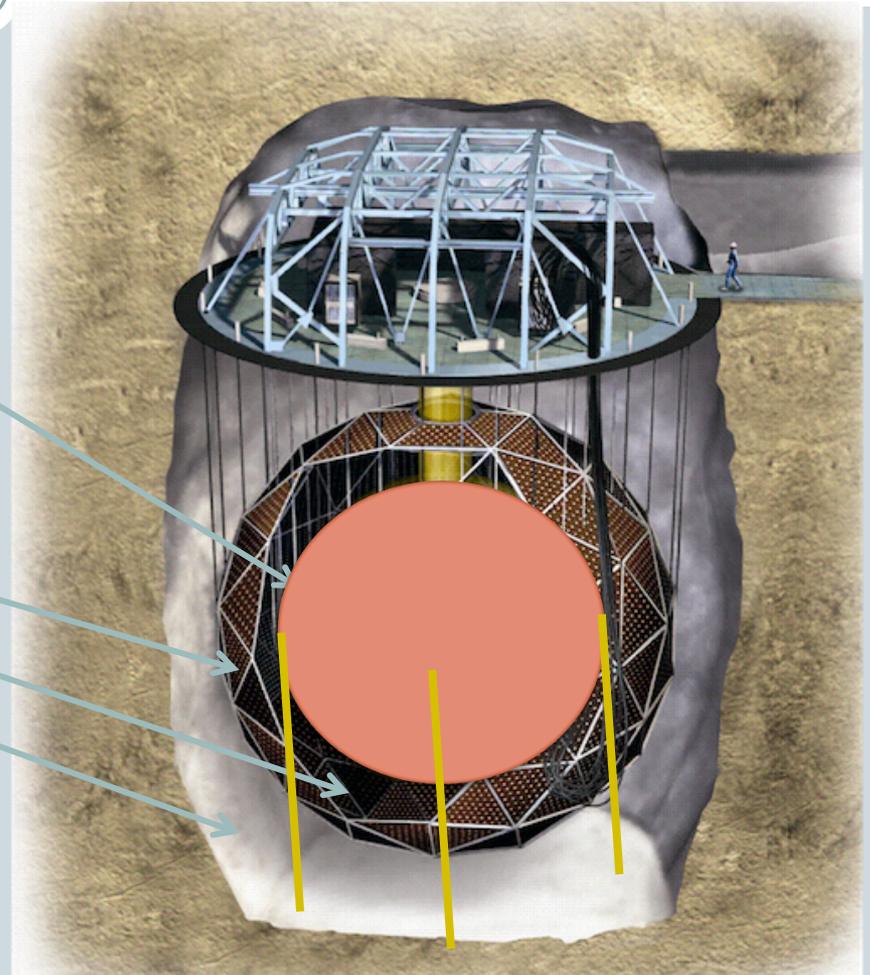


Does not include errors in the model other than scaling the mantle and crust

# SNO+ detector

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- Located 2 km underground ( $\sim 70$  muons/day) in the Vale Inco Ltd. Creighton Mine near Sudbury, Canada
  - 0.8 kton scintillator held in 12 m diameter acrylic vessel
  - 18 m diameter support structure holds 9500 PMTs ( $\sim 60\%$  photocathode coverage)
  - 1.7 kton inner shielding  $H_2O$
  - 5.3 kttons outer shielding  $H_2O$



# Future experiments

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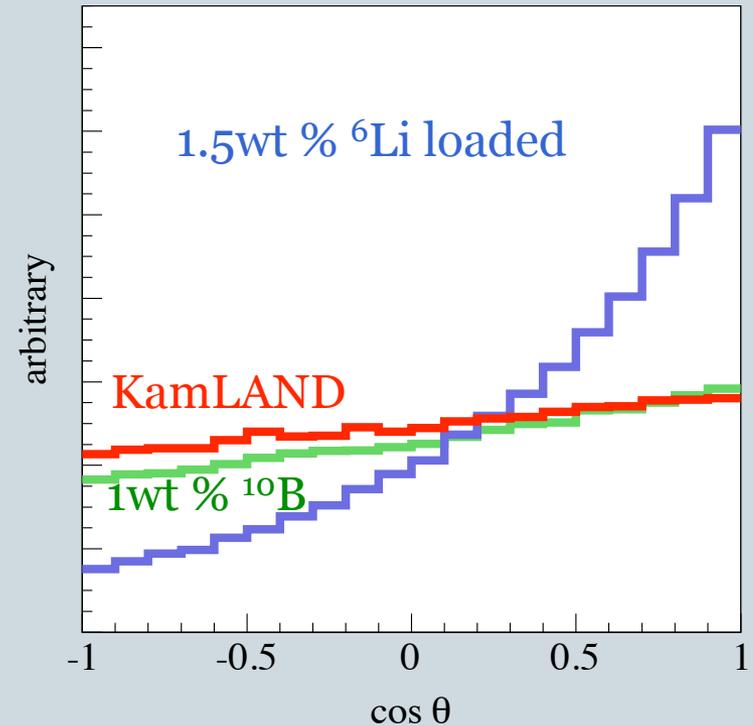
- SNO+ (0.8 kton), currently under construction
- Proposed detectors:
  - LENA (~50 kton), on continental crust
  - HanoHano (~10 kton), on oceanic crust
  - Baksan (~1 kton), on continental crust
  - DUSEL (??), on continental crust
  - EARTH (~1 kton), on oceanic or continental crust



# Directionality and $^4\text{K}$ detection

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- Directionality sensitive detectors could directly determine the U and Th distribution within Earth.
- K detection is also important. Since K is more volatile than U and Th, the absolute amount is less well known. Is there K in the core?

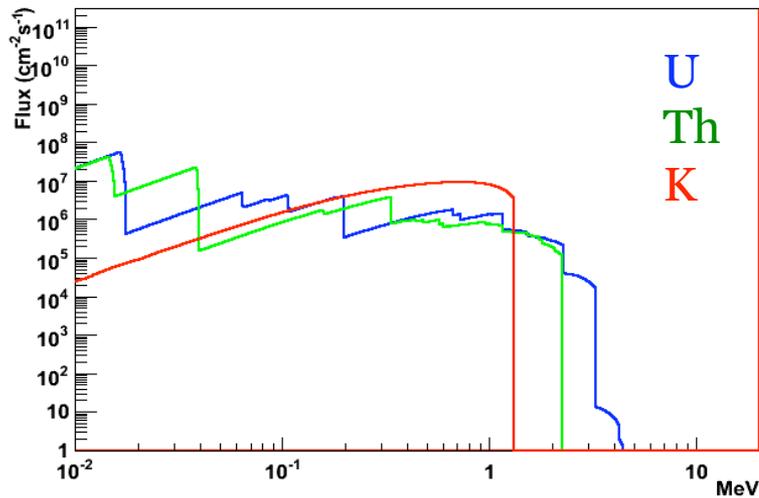


I. Shimizu. Nuclear Physics B  
(Proc. Suppl.) **168** (2007) 147

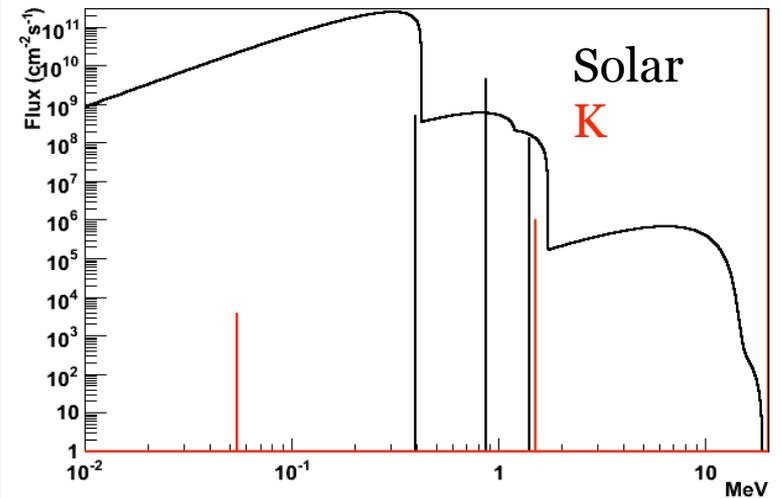
# Solar and geo neutrino signals

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## Antineutrinos



## Neutrinos



# Detection requirements

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- **Need to**
  - distinguish electron-antineutrinos from solar electron-neutrinos
  - be sensitive to neutrinos with energies  $< 3.4$  MeV geo-neutrinos
  - be sensitive to rates as low as  $\sim 20$  events/kton/year

# Other detectors

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## Liquid Ar

- Elastic scattering:
  - Low cross-section
  - Overwhelmed by solar neutrinos, therefore need good angular resolution
- CC,  $\nu_e + {}^{40}\text{Ar} \rightarrow e^+ + {}^{40}\text{Cl}^*$ :
  - Threshold too high

## Water Cherenkov

- Detect electron-antineutrinos from neutron inverse beta decay
$$\bar{\nu}_e + p \rightarrow e^+ + n$$
- Not sensitive to geo-neutrinos because they are below the Cherenkov threshold

# Conclusions

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- KamLAND and now Borexino have observed geo-neutrinos.
- The reported error is now approaching the level where they could contribute to our knowledge of Earth.
- Need to build more detectors to determine contribution from the mantle in order to help solve problem of mantle convection.