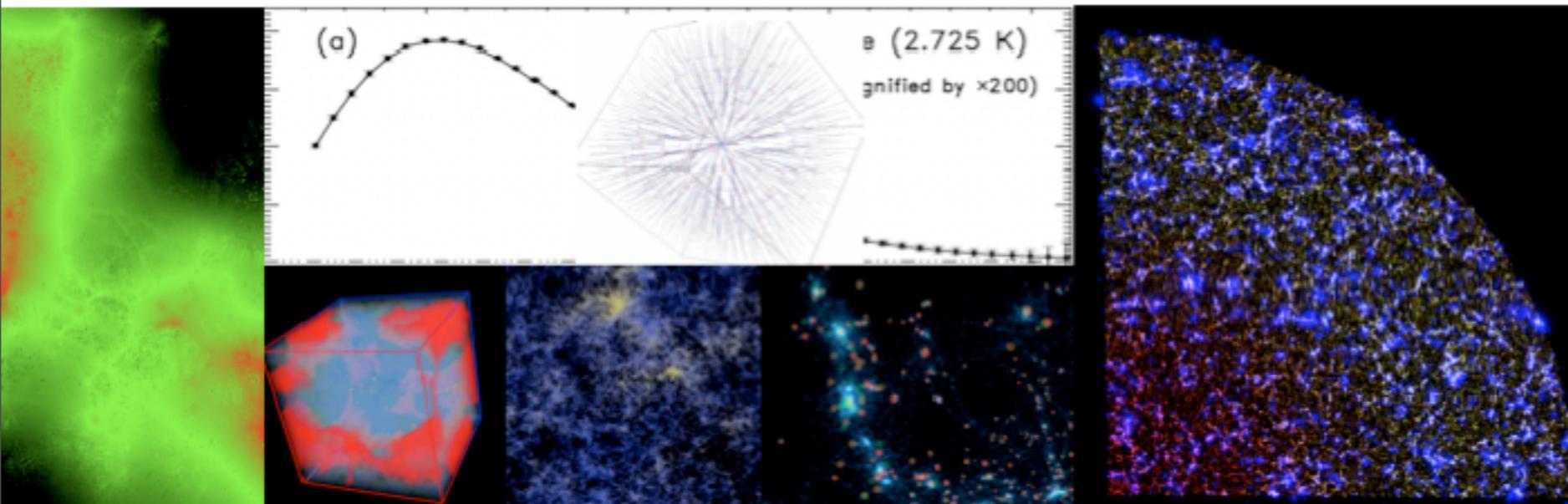


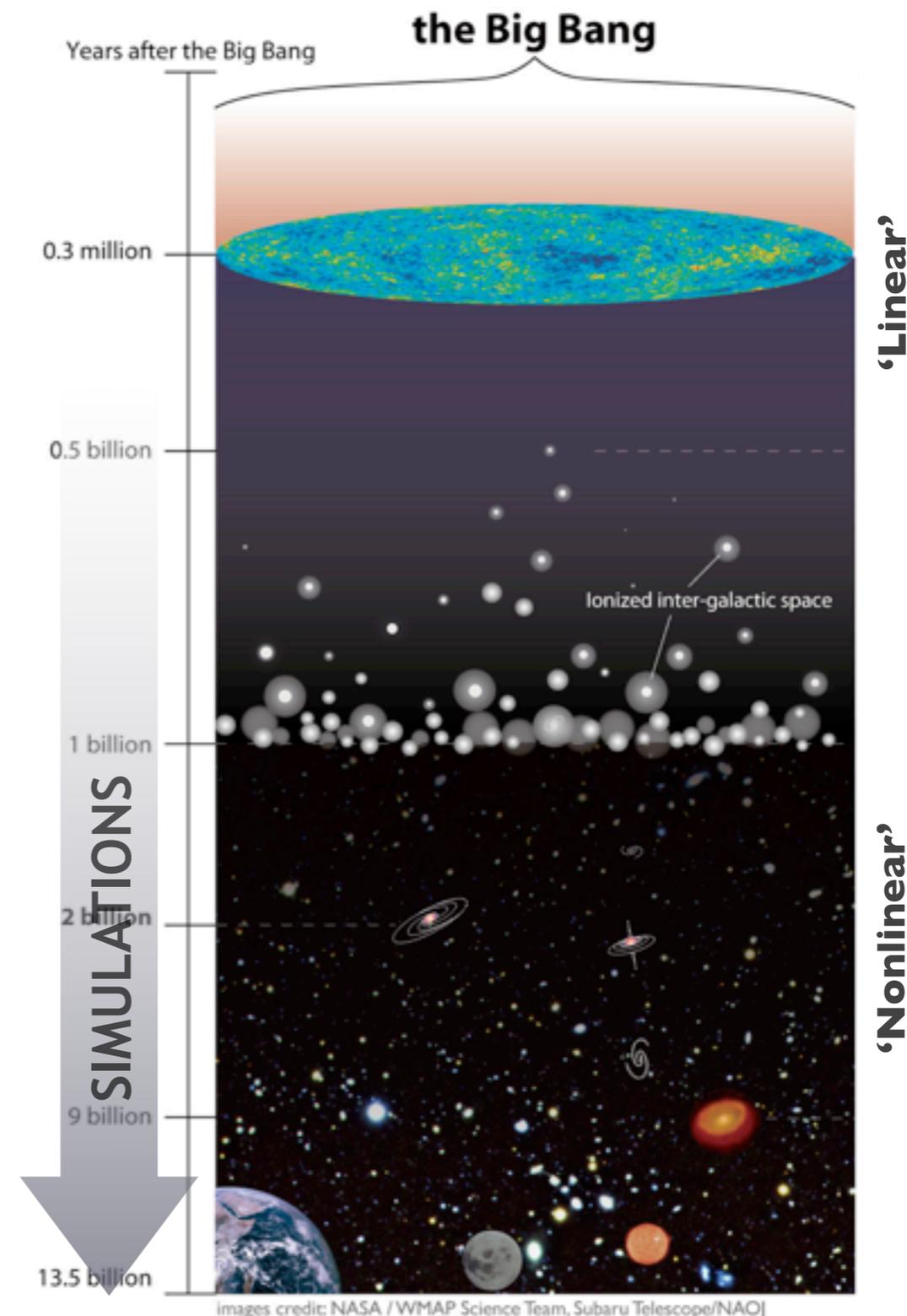
Summary/Survey: Computational Cosmology

Salman Habib



Structure Formation: The Basic Paradigm

- **Solid understanding of structure formation, success underpins most cosmic discovery**
 - Initial conditions laid down by inflation
 - Initial perturbations amplified by gravitational instability in a dark matter-dominated Universe
 - Relevant theory is gravity and atomic physics ('first principles')
- **Early Universe: Linear** perturbation theory very successful (CMB)
- **Latter half of the history of the Universe: Nonlinear** domain of structure formation, **impossible** to treat without large-scale computing



Prehistory: An Early Simulation

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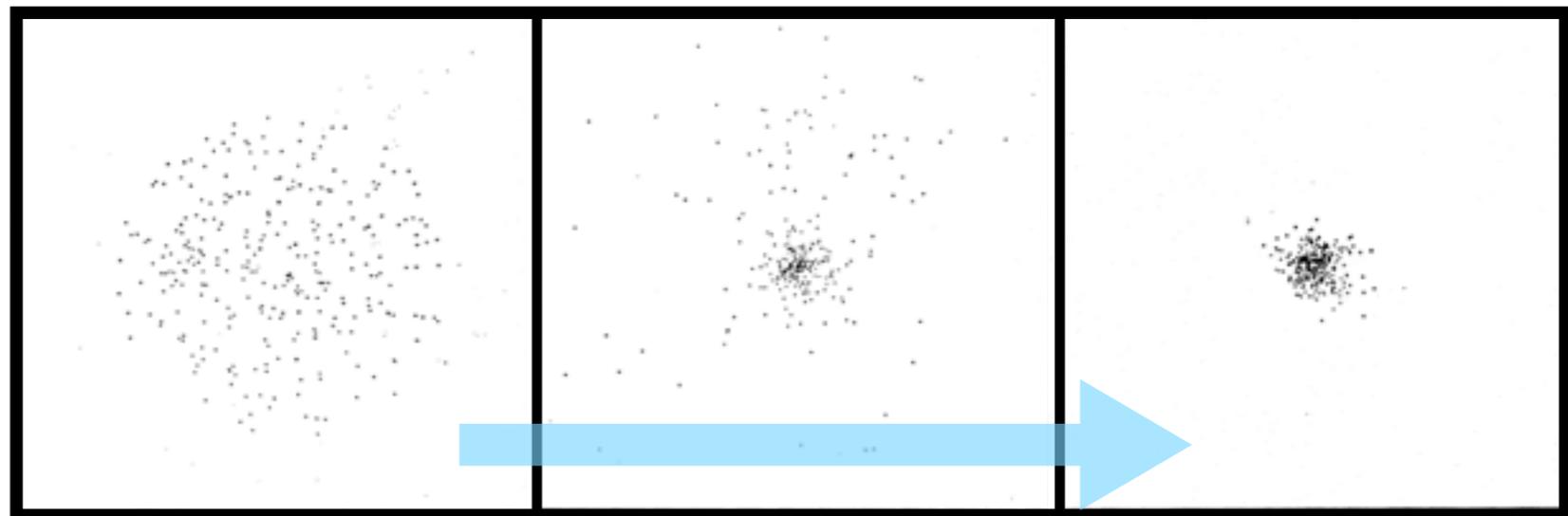
Structure of the Coma Cluster of Galaxies*

P. J. E. PEEBLES†

Palmer Physical Laboratory, Princeton University, Princeton, New Jersey

(Received 7 October 1969)

In some cosmologies, a cluster of galaxies is imagined to be a gravitationally bound system which, in analogy with the formation of the Galaxy, originated as a collapsing protocluster. It is shown that a numerical model based on this picture is consistent with the observed features of the Coma Cluster of galaxies. The cluster mass derived from this model agrees with previous values; however, an analysis of the observational uncertainty within the framework of the model shows that the derived mass could be consistent with the estimated total mass provided by the galaxies in the cluster.



- Suite of 300 (and less) particle simulations
- Run on a CDC 3600, ~1Mflops, 32KB+ at LANL
- Is **nine** orders of magnitude improvement in **both** performance and memory good enough for precision cosmology?

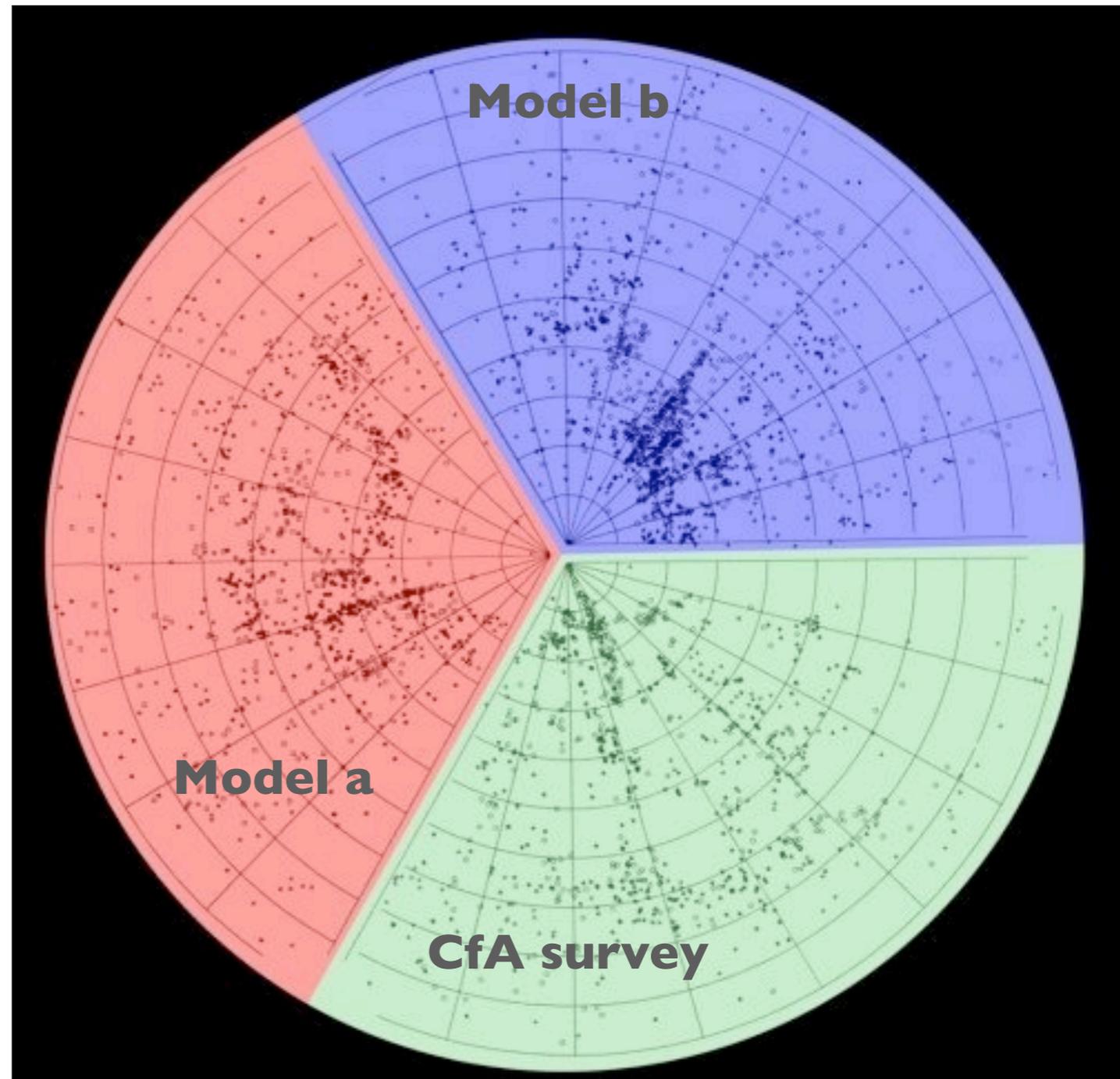


“The Universe is far too complicated a structure to be studied deductively, starting from initial conditions and solving the equations of motion.”

Robert Dicke (Jayne Lectures, 1969)

Fast Forward: A Compressed History

- **Beginnings (60s/70s)**
 - Primarily direct particle-particle methods, no theory of initial conditions
- **'Medieval' Period (80s)**
 - Inflation appears; Zeldovich approximation applied to ICs, application of state of the art simulation methods
- **'Modern' Period (90s)**
 - Multi-resolution parallel codes, gasdynamics simulations approach maturity
- **Current Era**
 - Transition from qualitative to quantitative predictions, driven by observations

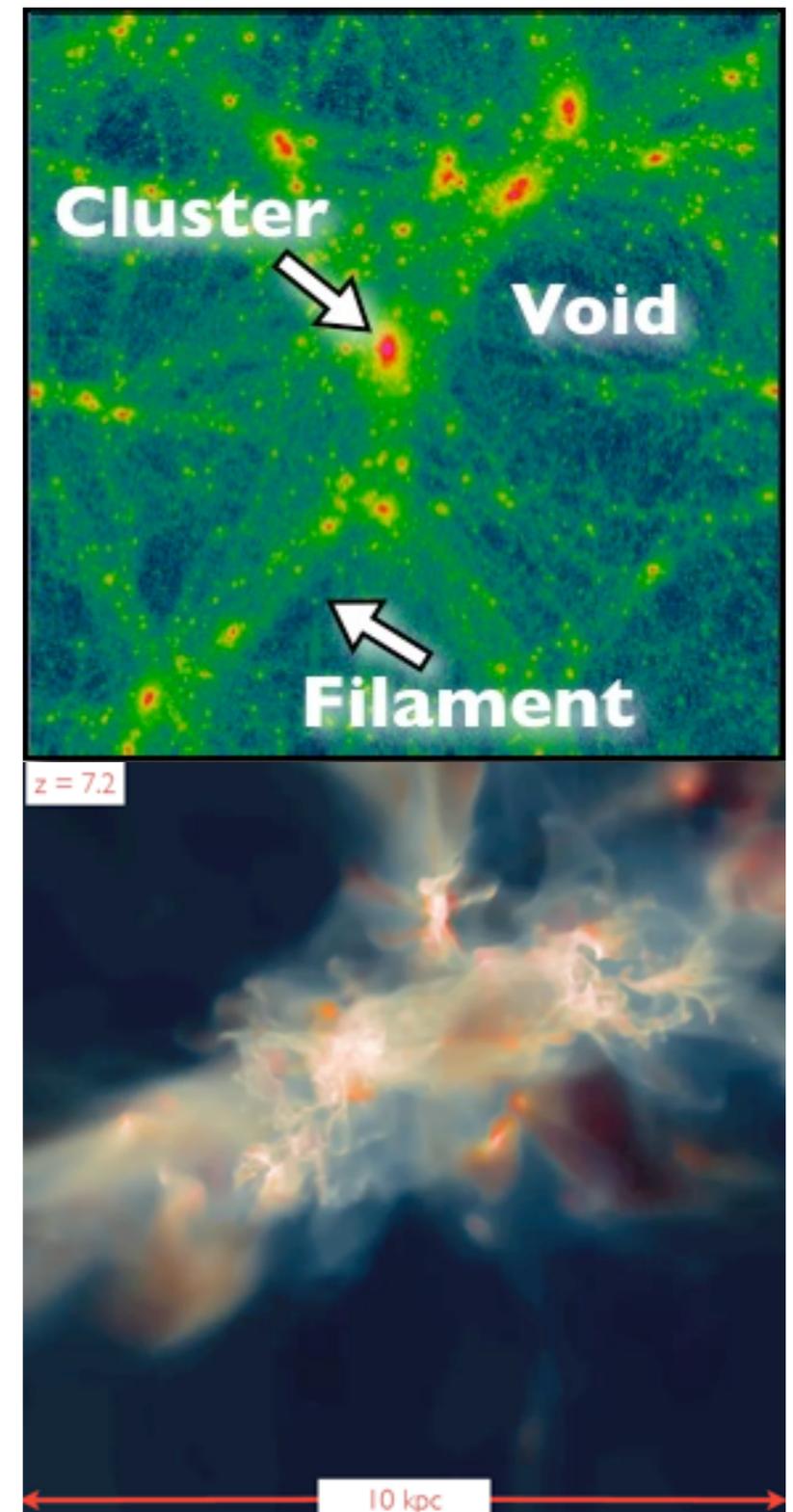


Davis et al 1985

Mock catalogs from 25 years ago for “eyeball” comparisons with the CfA galaxy survey

Achievements

- **Paradigm of Structure Formation**
 - “Cosmic Web” unifies a wide array of observations, from galaxy surveys to the Ly-alpha forest
 - Predictions for halo shapes, substructure, spatial statistics, halo bias (clustering relative to underlying density field), and their evolution
- **Cosmic Probes**
 - Halo abundance and mass distribution
 - Nonlinear domains of clustering probes
- **Physics of the IGM, clusters, reionization --**
 - Multi-resolution, multi-physics simulations essential tools for building more detailed picture, understanding of empirical relations and their scatter



Examples of Current Results

- **Cosmology Simulation Suites**

- LasDamas -- large suite for a single cosmology, many realizations and box sizes (SDSS mocks)
- Coyote Universe -- large simulation for multiple cosmologies, very fast ('instantaneous') emulation for observables

- **Galaxy Modeling for Surveys**

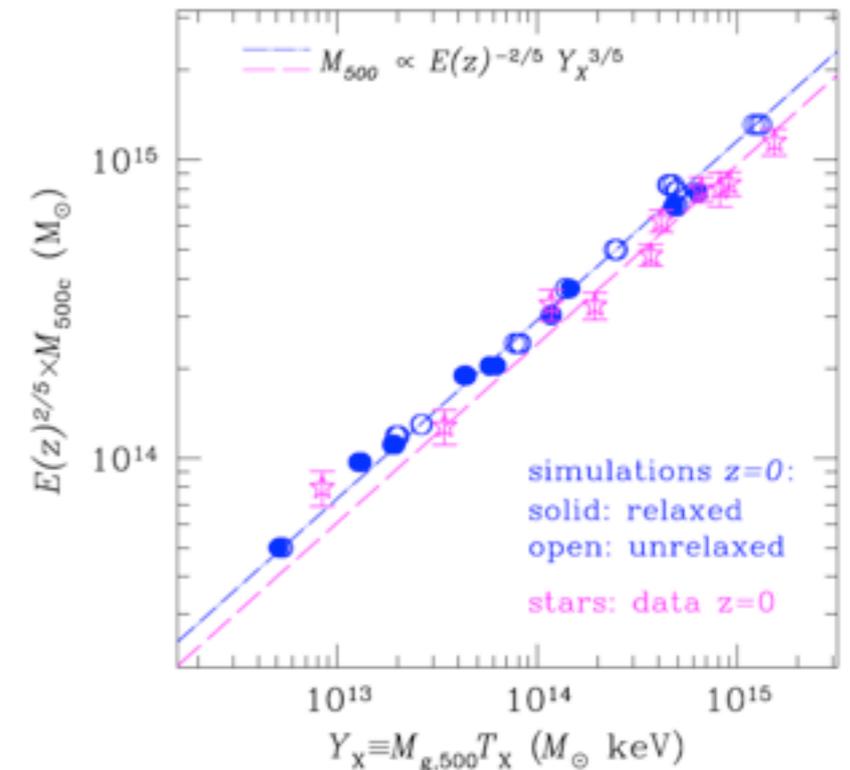
- High-resolution, 'small' volume simulation to follow details of structure formation (Bolshoi, Millennium II, MW TreePM, --)
- Baryonic effects on weak lensing

- **Cluster Physics**

- Realistic cluster simulations for understanding scatter of mass-observable relations

- **Ultra-high Resolution 'Milky Way'**

- Aquarius, Via Lactea, Silver River, --



Precision Cosmology: Modeling and Observations

- **Early Period**

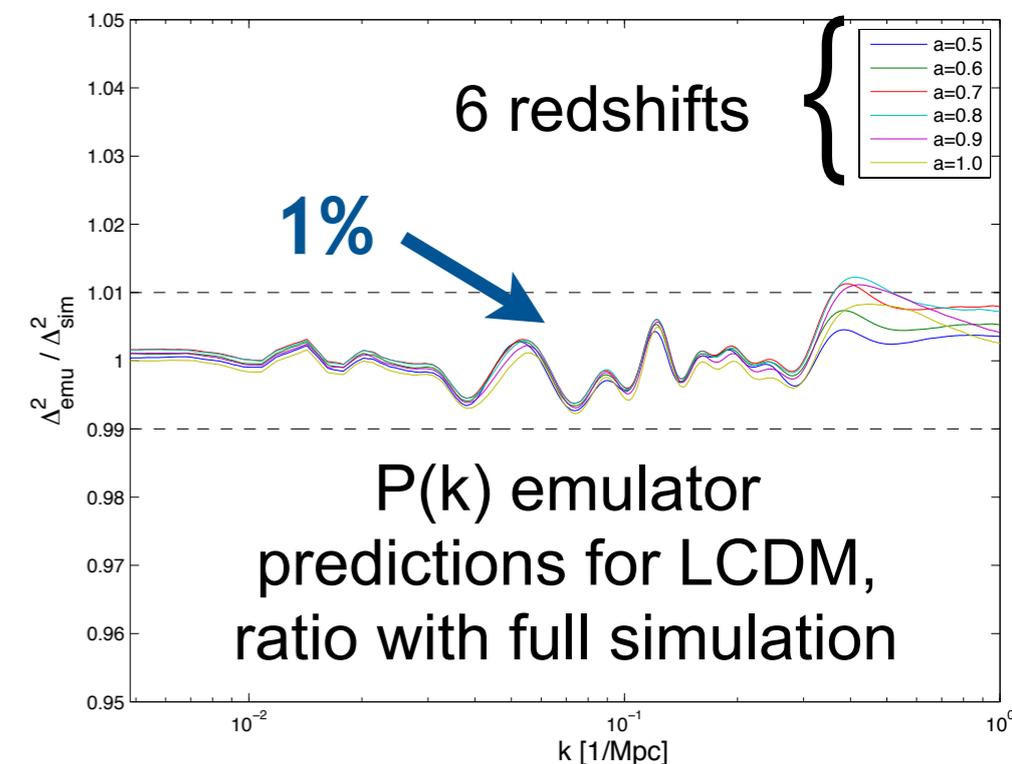
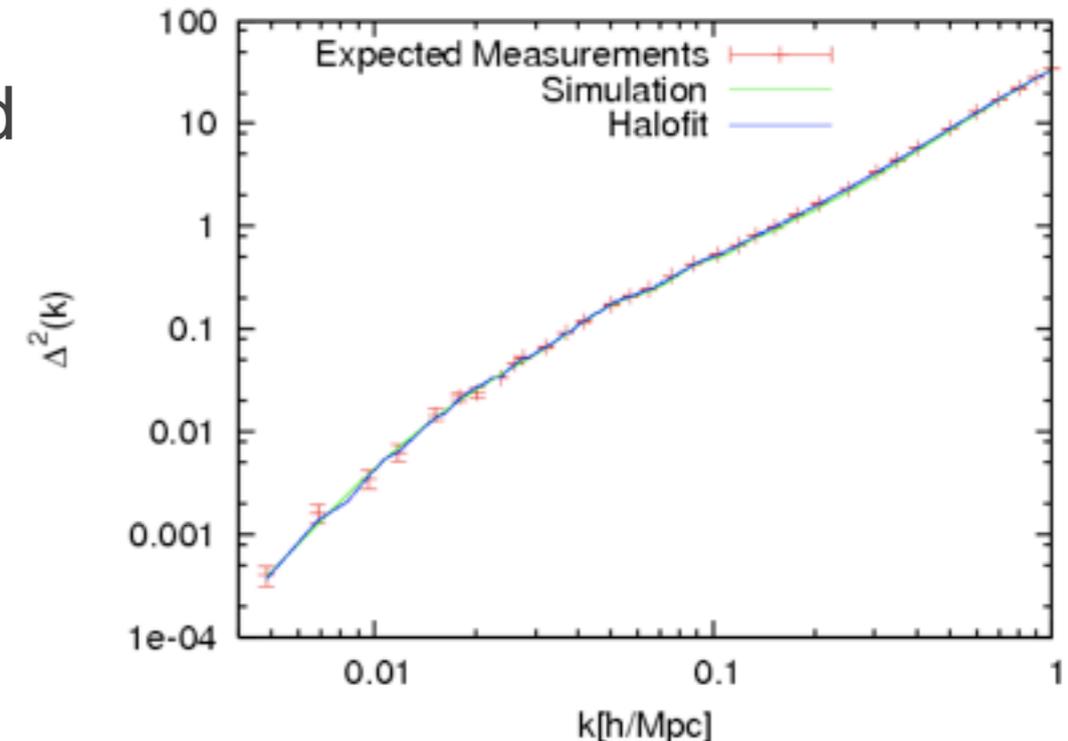
- Access to large data choked (technology did not exist), insignificant computing
- Characterized by small datasets, ‘eyeball’ comparisons, simple statistics

- **Intermediate Phenomenology (~10%)**

- Use simulation to build intermediate, simplified theoretical model; use this to interact with observations (HOD models, HaloFit, scaling relations from simulations, --)

- **‘Direct’ Numerical Phenomenology (~1%)**

- ‘Theory’ = direct interaction with observations by sophisticated simulations (or intermediate numerical products, ‘emulators’); complex space -- understand systematic errors (missing/wrong physics), bias



Precision Cosmology: Modeling and Observations

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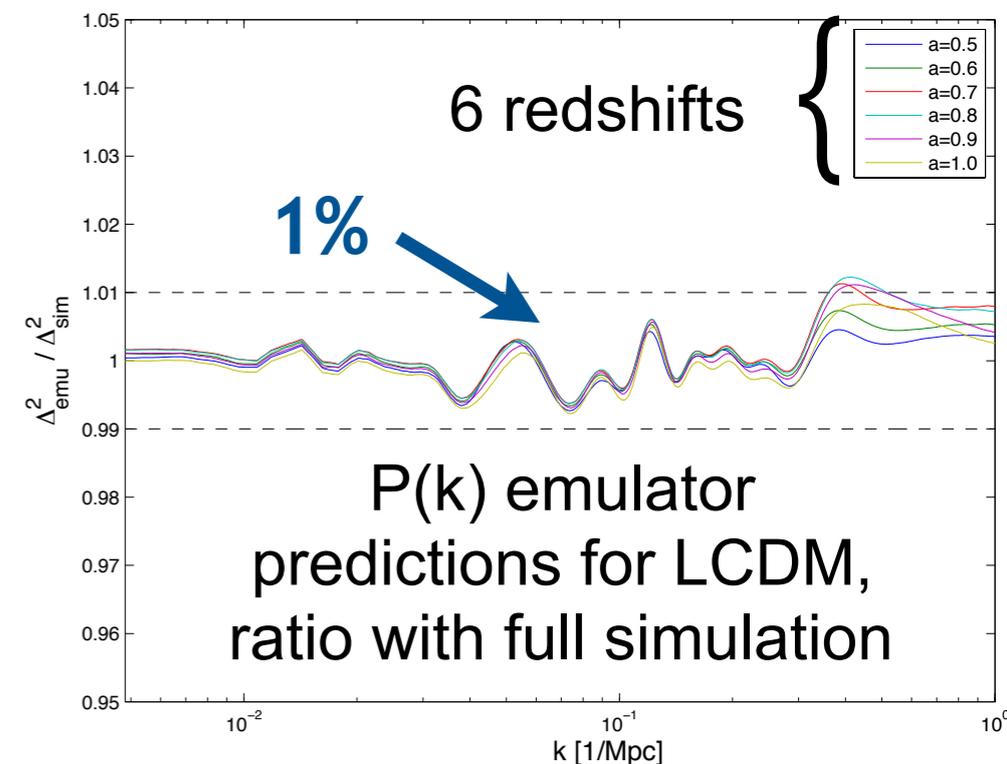
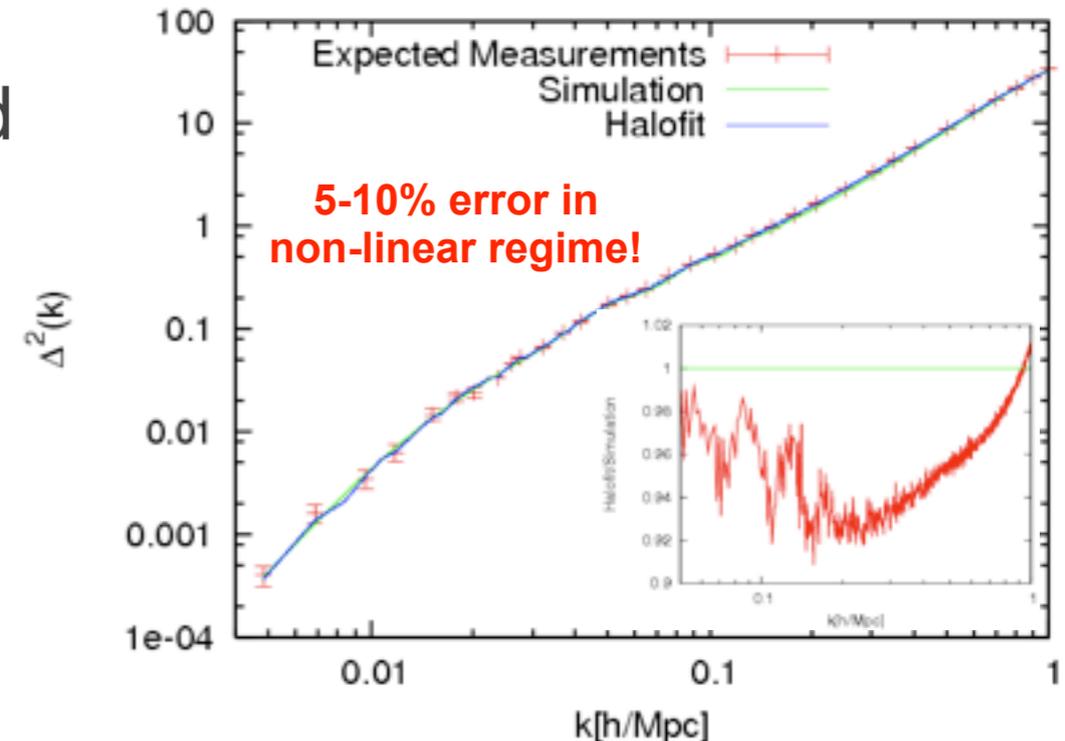
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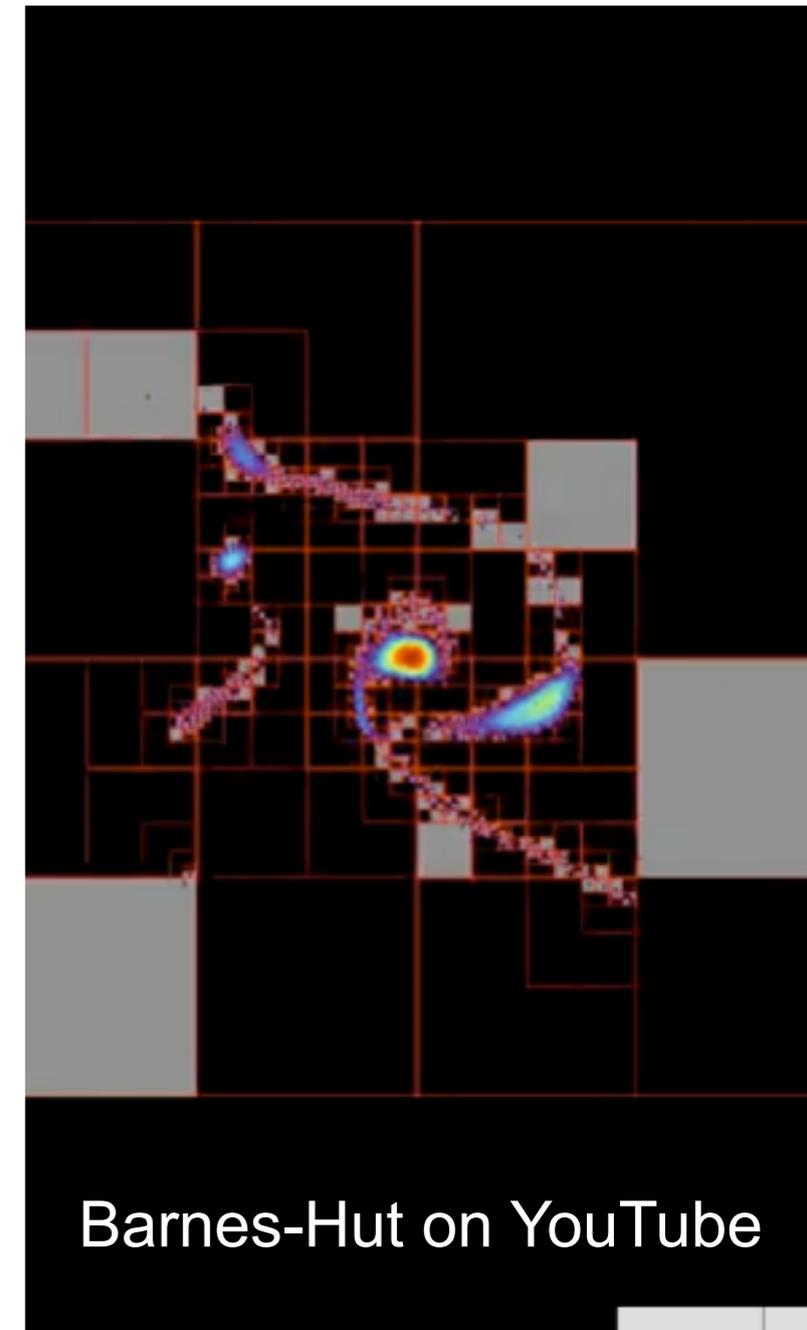
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Computational Cosmology and HPC

- **Early Era**
 - Early use of vector systems, methods borrowed from plasma physics (PM, adaptive P3M, AMR) plus tree algorithms
- **Parallel Simulations**
 - Early parallelization of all codes, efficient utilization of supercomputing resources (several Gordon Bell prizes, including custom hardware)
- **More Physics: Multi-Physics/Multi-Resolution**
 - Several cosmology codes have complexity similar to combustion codes, others are aimed at optimizing performance and memory (survey scale simulations)
- **Next-Generation Architectures**
 - Increased concurrency with communication bottlenecks and costs -- codes must evolve to deal with these issues



Computational Cosmology and Large Data

- **Digital Data**
 - Plate digitization
 - CCD sky surveys, SDSS (10s of TB), many 'use' modes -- very rich scientifically
- **Next-Generation Surveys**
 - DES starting (~PB total)
 - LSST 'threatening' (~100 PB total)
 - Planck ongoing (analysis at NERSC)
- **Analysis modes**
 - Many 'single' applications (each complex)
 - Very many cross-correlations to constrain systematics, tease out new signals
- **Problems of Scale**
 - Simulation data and observational data sizes enormous, need to implement new algorithms and methodologies

Help

Welcome to SkyServer! This site gives you access to the Digital Sky Survey (SDSS). Here, you will find information that professional astronomers use.

You are now viewing data from the Sloan Digital Sky Survey (SDSS) (opens in a new window)

Use the links below, or the links to the left,

Start Here

Site News lists the latest changes to the information

Introduction to SkyServer gives you a quick overview of how to use the data

Cooking with Sloan consists of guides for working with the data

FAQ is frequently asked questions about the data

Guide to Searching for Data

Search Form User Guide provides help on using the search form

SQL tutorial is an introduction to Structured Query Language (SQL) uses, and its syntax

Using SQL with SkyServer contains more advanced advice making your queries run fast

Sample SQL Queries are many real-life examples of queries

Graphing and Analyzing Data shows you how to do scientific analysis

Query Limits



U.S. DEPARTMENT OF
ENERGY

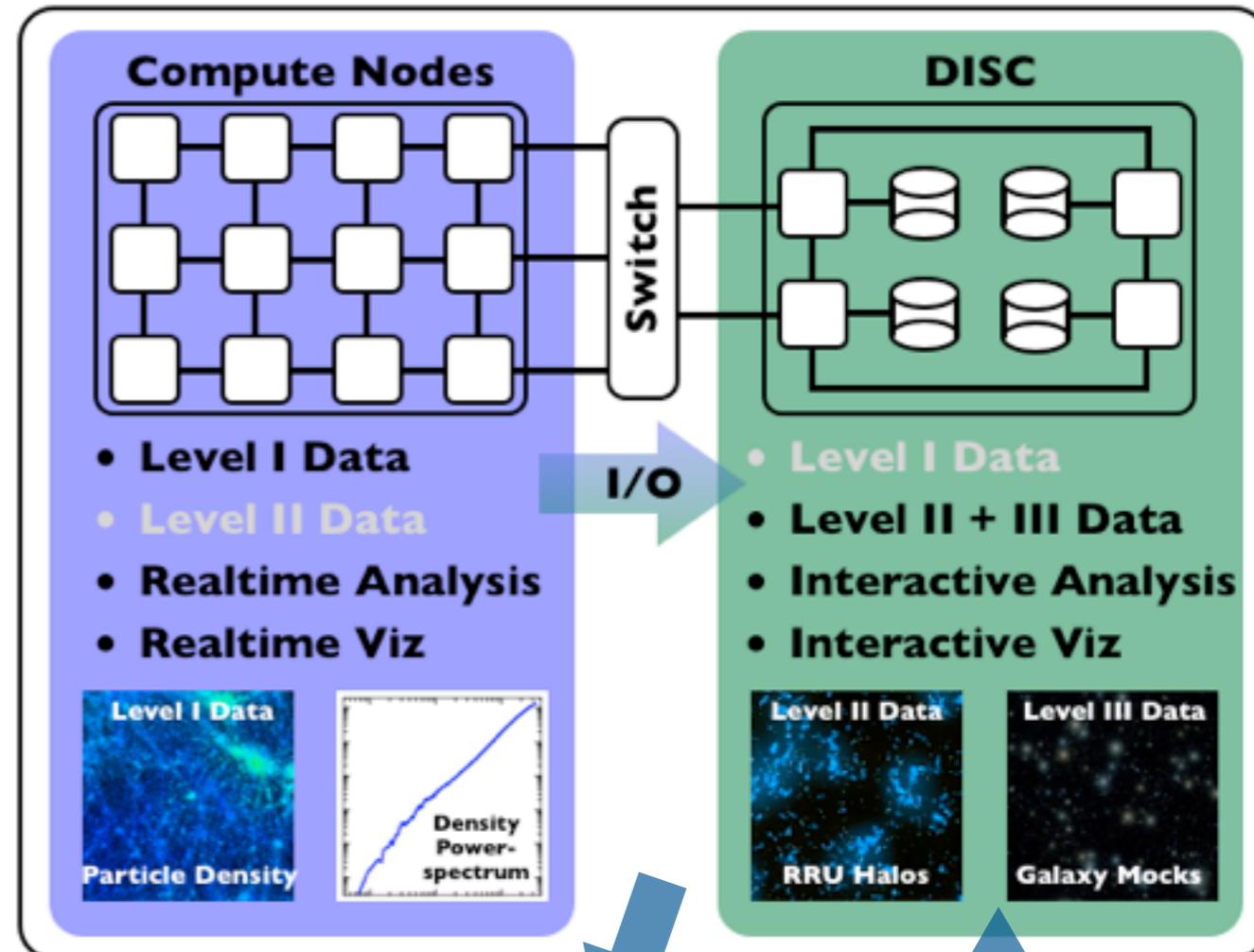
Office of
Science

'Ultimate Vision'

- **Clean up 'Discovery Space':** Robust theory with subtle signals
- **Precision Cosmic Calibration at Scale:** 'All Sky' solution of the cosmic inverse problems in the nonlinear regime
- **Cosmology Simulations at the Exascale:** Next-generation computing and beyond as essential theoretical and analysis tools



- **Large Data:** Simulation & observational datastreams: Archiving, serving, quality assurance, (joint) analyses
- **Simulation/Data/Analysis:** Cross-platform, multi-source, analysis and interrogation frameworks



Science Results!



Observation Feed

